



## SOLUTION

### TQ12 [15 points]

12.1 For a satellite that describes a uniform circular motion of radius  $r$  around the star and mass  $M$ , we have:

$$v = [R_{Hodograph}] = \sqrt{\frac{GM}{r}} \quad [1 \text{ point}]$$

12.2  $\vec{a} = -\frac{GM}{r^2} \hat{r}$  [1 point]

$$L = mr^2\omega \quad [1 \text{ point}]$$

$$\rightarrow \vec{a} = -\frac{GMm\omega}{L} \hat{r} = \frac{\Delta v}{\Delta t}$$

$$\frac{GMm}{L} \frac{\Delta\theta}{\Delta t} = \left| \frac{\Delta v}{\Delta t} \right|$$

$$\Delta v = \pm \frac{GMm}{L} \Delta\theta$$

$$k = \frac{GMm}{L} \quad [2 \text{ points}]$$

12.3  $L = mv_c R$  [1 point]

$$k_c = \frac{GMm}{L} = v_c = \sqrt{\frac{GM}{R}} \quad [1 \text{ point}]$$

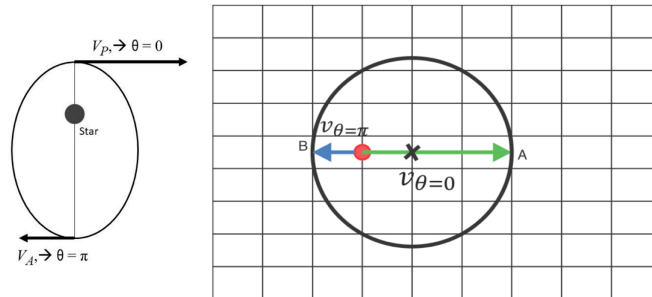
12.4 In the following scheme, the hodographic circumference has been constructed as follows:

- For  $\theta = 0$ , the velocity vector is drawn in arbitrary units (3, for instance) corresponding to the velocity in the periastron,  $v_p$ . Its head determines the point  $A$ .
- For  $\theta = \pi$ , the velocity vector is drawn diametrically opposite in the apastron,  $v_A$ , also in arbitrary units. Its head determines point  $B$ . The hodograph must pass through points  $A$  and  $B$ . Therefore, the radius of the hodograph must be equal to:

$$R_{Hodograph} = \frac{v_p + v_A}{2} \quad [1 \text{ point}]$$

And the "distance"  $d$  from the star to the center of the circumference will be:

$$d = \frac{v_p - v_A}{2} \quad [1 \text{ point}]$$



[2 points]

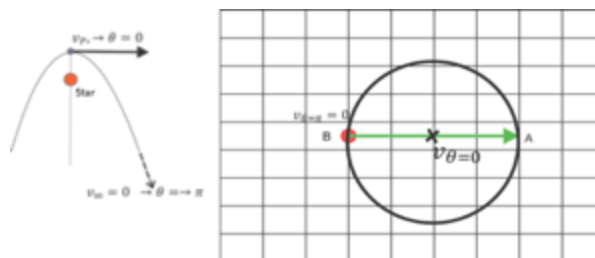
12.5 In the following scheme, the hodographic circumference has been constructed as follows:

- For  $\theta = 0$ , the velocity vector is drawn in arbitrary units (4, for instance) corresponding to the velocity in the periastron, that is  $v_p$ . Its head determines the point  $A$ . Then the velocity vector is drawn diametrically opposite in the apastron, that is,  $v_A = 0$ . Its head determines point  $B$  that coincides with the star. The hodograph must pass through points  $A$  and  $B$ . Therefore, the radius of the hodograph must be equal to:

$$R_{Hodograph} = \frac{v_p}{2} \quad [1 \text{ point}]$$

And the distance  $d$  from the star to the center of the hodographic circumference

$$d = \frac{v_p}{2} \quad [1 \text{ point}]$$



[2 points]